# Inverse Scattering Method Solves the Problem of Full Statistics of Nonstationary Heat Transfer in the Kipnis-Marchioro-Presutti Model 

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(Received 3 December 2021; revised 1 February 2022; accepted 8 March 2022; published 30 March 2022)


#### Abstract

We determine the full statistics of nonstationary heat transfer in the Kipnis-Marchioro-Presutti lattice gas model at long times by uncovering and exploiting complete integrability of the underlying equations of the macroscopic fluctuation theory. These equations are closely related to the derivative nonlinear Schrödinger equation (DNLS), and we solve them by the Zakharov-Shabat inverse scattering method (ISM) adapted by D. J. Kaup and A. C. Newell, J. Math. Phys. 19, 798 (1978) for the DNLS. We obtain explicit results for the exact large deviation function of the transferred heat for an initially localized heat pulse, where we uncover a nontrivial symmetry relation.


DOI: 10.1103/PhysRevLett.128.130602

Introduction.-Full statistics of currents of matter or energy in macroscopic systems away from thermodynamic equilibrium is a fundamental quantity that has attracted much attention from statistical physicists in the past two decades. Major progress has been achieved in determining this quantity for nonequilibrium steady states in simple models of interacting particles [1-4]. Nonstationary fluctuations of current, however, proved to be much harder for analysis [5-11].

A convenient and widely used family of models for studying the full statistics of currents is stochastic lattice gases [12-15]. One important example is the Kipnis-Marchioro-Presutti (KMP) model of heat transfer. The KMP model involves immobile particles occupying a whole lattice and carrying continuous amounts of energy. At each random move the total energy of a randomly chosen pair of nearest neighbors is randomly redistributed among them according to uniform distribution. The KMP model originally attracted much interest as the first model for which Fourier's law of heat diffusion at a coarse-grained level was proven rigorously [16]. By now it has become a paradigmatic model of nonequilibrium fluctuations of transport [4,6-8,11,17-27].

Here we study a full nonstationary heat-transfer statistics in the KMP model on an infinite one-dimensional lattice. Suppose that only one particle has a nonzero energy at $t=0$. Because of the energy exchange with the neighbors, the energy will start spreading throughout the system. At times much longer than the inverse rate of the energy exchange between the two neighbors (equal to $1 / 2$ ), and at distances much larger than the lattice constant (equal to 1 ), the mean coarse-grained temperature $\bar{u}(x, t)$ in the KMP
model is governed by the heat diffusion equation [12,14,16] $\partial_{t} \bar{u}(x, t)=\partial_{x}^{2} \bar{u}(x, t)$. The initial temperature is a delta function, $\bar{u}(x, t=0)=W \delta(x)$, and so the solution is

$$
\begin{equation*}
\bar{u}(x, t)=(W / \sqrt{4 \pi t}) \exp \left(-x^{2} / 4 t\right) \tag{1}
\end{equation*}
$$

However, in stochastic realizations of the KMP model the coarse-grained temperature will fluctuate around the expected profile $\bar{u}(x, t)$, see Fig. 1. To characterize these nonstationary fluctuations, we will consider the total amount of heat $W_{>}$, observed on the right half line $x>0$ at time $t=T \gg 1$. The expected value of $W_{>}$is $W / 2$, and we will study the full time-dependent statistics of the heat excess,

FIG. 1. Monte Carlo simulation of the KMP model with $W=1$. Plotted is the simulated temperature profile $u$ as a function of $x$ at time $t=1.5 \times 10^{4}$ (bars), its spatial average over each 50 consecutive lattice sites (solid line) and the theoretical Gaussian profile (1) (dashed line).

Obviously $\mathcal{P}(J, T)$, the probability distribution of $J$ at time $T$, has a compact support $|J| \leq W / 2$.

Similar nonstationary large-deviation settings, but with a steplike initial condition for the particle density or temperature, have been recently studied for a whole family of diffusive lattice gases [6-10], of which the KMP model is an important particular case. The main working tool of these studies has been the macroscopic fluctuation theory (MFT) [28]: a weak-noise theory, whose starting point is fluctuational hydrodynamics ( FH ) [12,14,29]. The FH is a coarse-grained description of the lattice gas, which is accurate when the characteristic length scale of the problem (here the diffusion length $\sqrt{T}$ ) and the observation time $T$ are much larger than the lattice constant 1 and the inverse elemental rate $1 / 2$ of the energy exchange, respectively. For diffusive lattice gases with a single conservation law the FH has the form of a single macroscopic Langevin equation, which accounts for the fluctuational contribution to the heat or mass flux. For the KMP model the Langevin equation reads $[12,14]$

$$
\begin{equation*}
\partial_{t} u=\partial_{x}\left(\partial_{x} u+\sqrt{2} u \eta\right) \tag{2}
\end{equation*}
$$

where $u(x, t)$ is the temperature, and $\eta(x, t)$ is a delta-correlated Gaussian noise: $\langle\eta(x, t)\rangle=0$ and $\left\langle\eta(x, t) \eta\left(x^{\prime}, t^{\prime}\right)\right\rangle=\delta\left(x-x^{\prime}\right) \delta\left(t-t^{\prime}\right)$.

The MFT $[28,30]$ relies on a saddle-point evaluation of the path integral for the stochastic process, described by Eq. (2). The small parameter of the saddle-point evaluation is again $1 / \sqrt{T} \ll 1$ : long times correspond to a weak noise. The saddle-point evaluation of the path integral boils down to a minimization of the action functional [30], constrained by the specified heat excess $J$ at $t=T$ and obeying the specified initial condition $u(x, t=0)$. For the statistics of the heat (or mass) excess, the MFT equations and boundary conditions in time were derived in Ref. [6], and we will present them shortly. For completeness, we also present their derivation in the Supplemental Material [30]. The solution of the MFT problem describes the optimal path of the process: the most likely time history of the temperature field $u(x, t)$ which dominates the probability distribution $\mathcal{P}(J, T)$ that we are after. The MFT problem, however, has proven to be very hard to solve analytically, especially for quenched (that is, deterministic) initial conditions [31]. In particular, for the KMP model, only small- $J$ [7] and large- $J$ [8] asymptotes have been obtained until now (but for a steplike initial condition).

This Letter reports a major advance in this area of statistical mechanics. We present an exact solution to the heat excess statistics problem by uncovering and exploiting complete integrability of the underlying MFT equations. We obtain explicit results for an initially localized heat pulse, $u(x, t=0)=W \delta(x)$, for which we uncover a nontrivial time-reversal mirror symmetry. These are the first exact non-steady-state large-deviation results for the
statistics of current in a lattice gas of interacting particles for quenched initial conditions.

Formulation of the MFT problem [6,30].-Let us rescale $t, x$, and $u$ by $T, \sqrt{T}$, and $W / \sqrt{T}$, respectively. The optimal path we are after is described by two coupled Hamilton equations for the rescaled temperature field $u(x, t)$ and the conjugate "momentum density" field $p(x, t)$, which describes the optimal history of the noise $\eta(x, t)$, conditioned on the heat excess $J$.

It is convenient to introduce the (minus) gradient field $v(x, t)=-\partial_{x} p(x, t)$. In the variables $u$ and $v$, the MFT equations take the form $[6,8,30]$

$$
\begin{gather*}
\partial_{t} u=\partial_{x}\left(\partial_{x} u+2 u^{2} v\right)  \tag{3}\\
\partial_{t} v=\partial_{x}\left(-\partial_{x} v+2 u v^{2}\right) \tag{4}
\end{gather*}
$$

The rescaled initial condition is

$$
\begin{equation*}
u(x, t=0)=\delta(x) \tag{5}
\end{equation*}
$$

The condition on the heat excess at $t=T$ becomes

$$
\begin{equation*}
\int_{0}^{\infty} u(x, t=1) d x-\frac{1}{2}=j \equiv \frac{J}{W} \tag{6}
\end{equation*}
$$

The minimization of the action functional, which enters the constrained path integral, with respect to variations of $u(x, t)$ yields, aside from Eqs. (3) and (4), a second boundary condition in time [6],

$$
\begin{equation*}
v(x, t=1)=-\lambda \delta(x) \tag{7}
\end{equation*}
$$

where $\lambda$ plays the role of a Lagrange multiplier, to be ultimately fixed by the constraint (6).

Once $u(x, t)$ and $v(x, t)$ are found, one can calculate the rescaled action, which can be written as [6-8]

$$
\begin{equation*}
s=\int_{0}^{1} d t \int_{-\infty}^{\infty} d x u^{2} v^{2} \tag{8}
\end{equation*}
$$

The action yields the probability density $\mathcal{P}(J, T, W)$ up to a preexponent:

$$
\begin{equation*}
\ln \mathcal{P}(J, T, W) \simeq-\sqrt{T} s\left(\frac{J}{W}\right) \tag{9}
\end{equation*}
$$

Since $\sqrt{T} \gg 1$, Eq. (9) has a clear large-deviation structure, and the action $s$ plays the role of a rate function.

A crucial and previously unappreciated observation is that Eqs. (3) and (4) coincide with the derivative nonlinear Schrödinger (DNLS) equation in imaginary time and space [32]. The DNLS equation (with real time and space) describes propagation of nonlinear electromagnetic waves in plasmas and other media [33]. An initial-value problem
for the DNLS equation is completely integrable via the Zakharov-Shabat inverse scattering method (ISM) adapted by Kaup and Newell for the DNLS [33]. The MFT formulation presents a difficulty, however, as here one needs to solve a boundary-value problem in time, rather than an initial-value problem. Here we overcome this difficulty by (i) making use of a shortcut that allows one to determine the rate function $s(j)$ even without the knowledge of $u(x, t)$ and $v(x, t)$ for all $t$, and (ii) exploiting a previously unknown symmetry relation [34], specific to the initial condition (5):

$$
\begin{equation*}
v(x, t)=-\lambda u(-x, 1-t) \tag{10}
\end{equation*}
$$

Solution of the MFT problem.-Equations (3) and (4) belong to a class of integrable systems for which a Lax pair exists, i.e., as we explain below, the equations are equivalent
to the compatibility condition of a system of two linear differential equations. The latter system defines scattering amplitudes which depend on $u$ and $v$. The idea behind the approach that we shall use-the ISM-is to consider the time evolution of these scattering amplitudes, which turns out to be very simple, as shown below. By relating these scattering amplitudes, at $t=0$ and $t=1$, to the fields $u$ and $v$, the method will enable us to find the heat excess $j=j(\lambda)$ which suffices for the calculation of $s=s(j)$.

Adapting the derivation of Kaup and Newell [33] to imaginary time and space, we consider the linear system

$$
\left\{\begin{array}{l}
\partial_{x} \boldsymbol{\psi}(x, t, k)=U(x, t, k) \boldsymbol{\psi}(x, t, k)  \tag{11}\\
\partial_{t} \boldsymbol{\psi}(x, t, k)=V(x, t, k) \boldsymbol{\psi}(x, t, k)
\end{array}\right.
$$

where $\boldsymbol{\psi}(x, t, k)$ is a column vector of dimension 2 ,

$$
U(x, t, k)=\left(\begin{array}{cc}
-i k / 2 & -i v \sqrt{i k}  \tag{12}\\
-i u \sqrt{i k} & i k / 2
\end{array}\right), \quad V(x, t, k)=\left(\begin{array}{cc}
k^{2} / 2-i k u v & -i(\sqrt{i k})^{3} v+i \sqrt{i k} \partial_{x} v-i \sqrt{i k} 2 v^{2} u \\
-i(\sqrt{i k})^{3} u+i \sqrt{i k} \partial_{x} u-i \sqrt{i k} 2 u^{2} v & -k^{2} / 2+i k u v
\end{array}\right),
$$

and $k$ is a spectral parameter. As one can check, the compatibility condition $\partial_{t} \partial_{x} \boldsymbol{\psi}=\partial_{x} \partial_{t} \boldsymbol{\psi}$, which corresponds to

$$
\begin{equation*}
\partial_{t} U-\partial_{x} V+[U, V]=0 \tag{13}
\end{equation*}
$$

is indeed equivalent to Eqs. (3) and (4).
Let us define the matrix $\mathcal{T}(x, y, t, k)$ as the $x$ propagator of the system (11), namely, the solution to

$$
\begin{equation*}
\partial_{x} \mathcal{T}(x, y, t, k)=U(x, t, k) \mathcal{T}(x, y, t, k) \tag{14}
\end{equation*}
$$

with $\mathcal{T}(x, x, t, k)=I$ (the identity matrix). At $x \rightarrow \pm \infty$, where the fields $u(x, t)$ and $v(x, t)$ vanish, the matrix $U$ becomes very simple,

$$
U(x \rightarrow \pm \infty, t, k)=\left(\begin{array}{cc}
-i k / 2 & 0  \tag{15}\\
0 & i k / 2
\end{array}\right)
$$

Therefore, it is natural to define the full-space propagator $G(t, k)$ as follows:

$$
\begin{align*}
G(t, k)= & \lim _{\substack{x \rightarrow \infty \\
y \rightarrow-\infty}}\left(\begin{array}{cc}
e^{i k x / 2} & 0 \\
0 & e^{-i k x / 2}
\end{array}\right) \\
& \times \mathcal{T}(x, y, t, k)\left(\begin{array}{cc}
e^{-i k y / 2} & 0 \\
0 & e^{i k y / 2}
\end{array}\right) \tag{16}
\end{align*}
$$

The entries of the matrix $G(t, k)$ are the scattering amplitudes of the system (11). The time evolution of $G(t, k)$ is easy to find. Indeed, the matrix $\mathcal{T}(x, y, t, k)$ satisfies

$$
\begin{align*}
\partial_{t} \mathcal{T}(x, y, t, k)= & V(x, t, k) \mathcal{T}(x, y, t, k) \\
& -\mathcal{T}(x, y, t, k) V(y, t, k) \tag{17}
\end{align*}
$$

One can check that Eq. (17) is compatible with Eq. (14) (i.e., $\partial_{t} \partial_{x} \mathcal{T}=\partial_{x} \partial_{t} \mathcal{T}$ ) due to Eq. (13). The matrix $V(x, t, k)$ too becomes very simple in the limit $x \rightarrow \pm \infty$,

$$
V(x \rightarrow \pm \infty, t, k)=\frac{k^{2}}{2}\left(\begin{array}{cc}
1 & 0  \tag{18}\\
0 & -1
\end{array}\right)
$$

Plugging Eq. (18) into Eq. (17), one finds the time evolution of $\mathcal{T}(x \rightarrow \infty, y \rightarrow-\infty, t, k)$ which in turn, using Eq. (16), yields that of $G(t, k)$ :

$$
\begin{align*}
G(t, k) & =\left(\begin{array}{cc}
a(t, k) & \tilde{b}(t, k) \\
b(t, k) & \tilde{a}(t, k)
\end{array}\right) \\
& =\left(\begin{array}{cc}
a(0, k) & \tilde{b}(0, k) e^{k^{2} t} \\
b(0, k) e^{-k^{2} t} & \tilde{a}(0, k)
\end{array}\right) \tag{19}
\end{align*}
$$

where we have introduced here a notation for the matrix elements of $G(t, k)$.

Plugging the temporal boundary conditions (5) and (7), we calculate $G(0, k)$ and $G(1, k)$ explicitly by solving Eq. (14), see Ref. [30]. Comparing the two solutions and using Eq. (19) we obtain

$$
\begin{equation*}
i k\left[Q_{+}(k)+Q_{-}(k)\right]-i k Q_{-}(k) \times i k Q_{+}(k)=-\lambda i k e^{-k^{2}} \tag{20}
\end{equation*}
$$



FIG. 2. Analytical results for $Q_{+}(k)$, described by Eqs. (22), (23), and (25) (solid lines), versus numerical results (dashed lines) for $\lambda=1$, or $j=0.09568 \ldots$. The symmetric and antisymmetric curves show $\operatorname{Re} Q_{+}(k)$ and $\operatorname{Im} Q_{+}(k)$, respectively.
where $Q_{ \pm}(k)$ are the Fourier transforms of $v(z, 0)$ restricted to $z>0$ and $z<0$, respectively:

$$
\begin{align*}
& Q_{-}(k)=\int_{-\infty}^{0} v(z, 0) e^{i k z} d z \\
& Q_{+}(k)=\int_{0}^{\infty} v(z, 0) e^{i k z} d z \tag{21}
\end{align*}
$$

We solve Eq. (20) in Ref. [30], with the result

$$
\begin{gather*}
i k Q_{ \pm}(k)=1-\left(1 \pm v_{ \pm}\right) e^{\Phi_{ \pm}(k)},  \tag{22}\\
\Phi_{ \pm}(k)= \pm \int_{-\infty}^{\infty} \frac{\ln \left(1+i \lambda k^{\prime} e^{-k^{\prime 2}}\right)}{k^{\prime}-k \mp i 0^{+}} \frac{d k^{\prime}}{2 \pi i}, \tag{23}
\end{gather*}
$$

where $v_{ \pm}=v\left(0^{ \pm}, 0\right)$.
To compute $v_{ \pm}$, we demand that $Q_{ \pm}(k)$ be regular at the origin, corresponding to a vanishing $v(z, 0)$ at infinity. Setting $k=0$ in Eq. (22) and using the Sokhotski-Plemelj formula

$$
\begin{equation*}
\int_{-\infty}^{\infty} \frac{f(k)}{k \pm i 0^{+}} \frac{d k}{2 \pi i}=\int_{\infty}^{\infty} \frac{f(k)}{k} \frac{d k}{2 \pi i} \mp \frac{1}{2} f(0) \tag{24}
\end{equation*}
$$

we obtain after some algebra

$$
\begin{equation*}
\pm v_{ \pm}=\exp \left[\mp \int_{-\infty}^{\infty} \arctan \left(\lambda k^{\prime} e^{-k^{\prime 2}}\right) \frac{d k^{\prime}}{2 \pi k^{\prime}}\right]-1 \tag{25}
\end{equation*}
$$

Taking the derivative of Eq. (22) with respect to $k$ at $k=0$, yields [30]

$$
\begin{equation*}
Q_{+}(0)=\frac{1}{4 \pi} \int_{-\infty}^{\infty} \frac{\ln \left(1+\lambda^{2} k^{2} e^{-2 k^{2}}\right)}{k^{2}} d k-\frac{\lambda}{2} . \tag{26}
\end{equation*}
$$

Figure 2 shows $\operatorname{Re} Q_{+}(k)$ and $\operatorname{Im} Q_{+}(k)$ versus $k$ at $\lambda=1$, obtained by plugging Eq. (25) for $v_{ \pm}$into Eq. (22). This figure also shows the same quantities computed by solving Eqs. (3) and (4) numerically with a back-and-forth iteration algorithm [35]. The analytical and numerical curves are almost indistinguishable.


FIG. 3. The exact rate function $s(j)$, given by Eqs. (27) and (29) (solid line) and two asymptotes: $s(|j| \ll 1)=\sqrt{8 \pi} j^{2}$ and Eq. (31) (dashed lines). Symbols: properly rescaled data from $10^{6}$ direct Monte-Carlo simulations of the microscopic KMP model for $T=10^{2}$, see Ref. [30] for details.

Using Eqs. (6), (10), and (21) alongside with the conservation law $\int_{-\infty}^{\infty} u(x, t) d x=1$, we determine $j=j(\lambda)$ :
$j(\lambda)=\frac{Q_{+}(0)}{\lambda}+\frac{1}{2}=\frac{1}{4 \pi \lambda} \int_{-\infty}^{\infty} \frac{\ln \left(1+\lambda^{2} k^{2} e^{-2 k^{2}}\right)}{k^{2}} d k$.
Now we use a shortcut which makes the results we have obtained so far sufficient for obtaining the rate function $s=s(j)$. The shortcut comes in the form of the relation $d s / d j=\lambda$, which follows from the fact that $j$ and $\lambda$ are conjugate variables, see, e.g., Ref. [36]. It allows one to calculate $s(j)$ bypassing Eq. (8) [which would require the knowledge of the whole optimal path $u(x, t)$ ]. We have

$$
\begin{equation*}
\frac{d s}{d \lambda}=\frac{d s}{d j} \frac{d j}{d \lambda}=\lambda \frac{d j}{d \lambda}=\frac{d Q_{+}(0)}{d \lambda}-\frac{Q_{+}(0)}{\lambda} \tag{28}
\end{equation*}
$$

Using Eq. (26), we integrate Eq. (28) with respect to $\lambda$ to get

$$
\begin{equation*}
s(\lambda)=Q_{+}(0)+\int_{-\infty}^{\infty} \frac{\operatorname{Li}_{2}\left(-\lambda^{2} k^{2} e^{-2 k^{2}}\right)}{8 \pi k^{2}} d k+\frac{\lambda}{2} \tag{29}
\end{equation*}
$$

where $\operatorname{Li}_{2}(z)=\sum_{k=1}^{\infty} z^{k} / k^{2}$ is the dilogarithm function, $Q_{+}(0)$ is given by Eq. (26), and the integration constant was determined from $s(\lambda=0)=0$. Equations (27) and (29) give the complete rate function $s(j)$ in a parametric form and represent the main result of this work. The exact optimal history of the temperature profile $u(x, t)$ proved difficult to obtain analytically, but it can be computed numerically [30].

Figure 3 shows $s(j)$ alongside with two asymptotes: $j \rightarrow 0$ and $|j| \rightarrow 1 / 2$, which correspond to $\lambda \rightarrow 0$ and $|\lambda| \rightarrow \infty$, respectively. Also shown are results of Monte Carlo simulations. The asymptote $\lambda \rightarrow 0$ can be obtained either from the exact rate function (27) and (29) [30], or from a perturbative expansion applied directly to the MFT equations [7]. By virtue of the symmetry (10), the latter can be done very easily. Indeed, in the leading order in $\lambda \ll 1$ Eqs. (8), (10), and (1) yield
$s(\lambda) \simeq \lambda^{2} \int_{0}^{1} d t \int_{-\infty}^{\infty} d x \bar{u}^{2}(x, t) \bar{u}^{2}(-x, 1-t)=\frac{\lambda^{2}}{8 \sqrt{2 \pi}}$.

The shortcut relation $d s / d j=\lambda$ can be rewritten as $(d s / d \lambda)(d \lambda / d j)=\lambda$. Combined with Eq. (30) it yields $s(j \rightarrow 0) \simeq \sqrt{8 \pi} j^{2}$. Then, from Eq. (9), we see that typical fluctuations of $J$ are normally distributed with variance $W^{2} /(32 \pi T)^{1 / 2}$. The $T^{-1 / 2}$ scaling of the variance should be contrasted with the $T^{1 / 2}$ scaling, obtained for a steplike initial condition [5-7]. However, the relative magnitude of the fluctuations-the ratio of the standard deviation and the average transferred heat-has the same scaling $T^{-1 / 4} \ll 1$ in both settings, as to be expected from the law of large numbers.

The asymptote of $|\lambda| \rightarrow \infty$ is more subtle [30]. The final result, already in terms of $j$, is

$$
\begin{align*}
s(|j| \rightarrow 1 / 2) & \simeq \frac{4\left[-\frac{1}{2} W_{-1}\left(-\frac{1}{2} \pi^{2} \Delta^{2}\right)\right]^{3 / 2}}{3 \pi} \\
& =\frac{4}{3 \pi} \ln ^{3 / 2}\left(\frac{2}{\pi \Delta} \sqrt{\left.\ln \frac{2}{\pi \Delta} \sqrt{\ln \frac{2}{\pi \Delta} \cdots}\right)},\right. \tag{31}
\end{align*}
$$

where $\Delta \equiv 1 / 2-|j| \ll 1$, and $W_{-1}(\ldots)$ is the proper branch of the product $\log$ (Lambert $W$ ) function [37]. At $j=1 / 2$, $s$ diverges and $\mathcal{P}$ vanishes, as to be expected. Nested-log large-current asymptotes similar to Eq. (31) appear to be typical for the KMP model and other models of the hyperbolic universality class $[8,9,11]$.

Discussion.-By combining the MFT and the ISM, we calculated exactly the rate function $s(j)$, see Eqs. (27) and (29), which describes the full long-time statistics of nonstationary heat transfer in the KMP model for an initially localized heat pulse. This is the first exact non-steady-state large-deviation result for the statistics of current in a lattice gas of interacting particles for quenched initial conditions. It opens the way to extensions of the ISM to additional fluctuating quantities of the KMP model. Another challenging goal is to apply the ISM to the simple symmetric exclusion process (SSEP) [12-15]-a lattice-gas model with quite different properties [9]. Encouragingly, the MFT equations for the SSEP (see, e.g., Ref. [6]) can be mapped to Eqs. (3) and (4) via a canonical transformation. This transformation, however, complicates the boundary conditions in time.

From a more general perspective, the MFT of lattice gases is a particular case of the weak-noise theory, or optimal fluctuation method (OFM): a highly versatile framework which captures a broad class of large deviations in macroscopic systems. For nonstationary processes the OFM equations-coupled nonlinear partial differential equations for the optimal path-are usually very hard to solve exactly. One class of problems of this type, which has received much recent attention, deals with the complete one-point height statistics of an interface whose dynamics is described by the Kardar-Parisi-Zhang equation [38]. The OFM captures the complete Kardar-Parisi-Zhang height statistics at short times [39-44]. Here too, a previous
analytical progress in the solution of the OFM equations was limited to asymptotics of very large or very small interface height. But very recently these OFM equationswhich coincide with the nonlinear Schrödinger equation (NLS) (not the derivative one) [42]-have been solved exactly $[45,46]$ by the ISM for several "standard" initial conditions. The two integrable systems, the NLS and DNLS, are closely related, so our approach can be compared with that of Refs. [45,46]. We used only standard techniques of the ISM which do not rely on additional tools, such as Fredholm determinants used in Refs. [45,46]. Because of its relative simplicity our approach appears to be more readily adaptable to solving additional largedeviation problems [47].

The research of E. B. and B. M. is supported by the Israel Science Foundation (Grants No. 1466/15 and No. 1499/20, respectively). E. B. is also supported by the US-Israel Binational Science Foundation, Grant No. 2020193. N. R. S. acknowledges support from the Yad Hanadiv fund (Rothschild fellowship).
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